Search for N*(1685) in Real Compton Scattering: First Results



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Why Compton Scattering?

Recent observation of N*(1685) resonance in eta photoproduction off the neutron dictates a need to search for its signal in other reactions. The decay of this resonance to a pion final state is predicted to be suppressed (R. A. Arndt, Ya. Azimov, M.v.Polyakov, I. Strakovsky, and R. Workman, Phys. Rev. C69 035208 (2004); arXiv: nuclth/0312126, M.Polyakov, A.Rathke, hep-ph/0303138)

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There two other reactions to search for N*(1685)

γn→K⁰Λ

The threshold is ~1.62 GeV. Problems are near-threshold effects overlapping with Fermi motion smearing, FSI, background reactions $\gamma N \rightarrow \pi \pi \pi N$, $\gamma N \rightarrow \pi \pi \pi N$, $\gamma N \rightarrow \omega N$, $\gamma N \rightarrow \rho N$ all having peculiarities near 1.72 GeV, plus low efficiency of the GRAAL detector for this final state .



GRAAL Setup



GRAAL forward lead-scintillator wall (``Russian Wall")

V.Kouznetsov et al., NIM A 487 (2002) 396.

An assembly of 16 modules. Each module is a sandwich of four 3000x40 mm2 bars with 3 mm thick lead plates between them. A 25 mm thick steel plate at the front of the module acts as a main converter and as a module support.



Particle identification and performance

√²⁰⁰ √⁹√ 150 Performance of the Russian Wall at GRAAL: TOF resolution -0.6 ns(FWHM) Angular resolution – 2-3 100 p deg(FWHM) 50 Photon efficiency – 95% Neutron efficiency – 22% 0 20 15 25 10 TOF, nsec a) √²⁰⁰ √e/ 150 100 n 50 0 10 20 30 40 50 b) TOF, nsec

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The GRAAL detector makes it possible the simultaneous study

үр→ηр	γn→ηn
γр→π ⁰ р	γn→π ⁰ n
үр→үр	γn→γn



on the quasi-free proton and quasi-free neutron in the same experimental run.

The main problem of Compton scattering measurements is the π^0 background.

Compton scattering

π^0 background



First Step: Study of π⁰ background SAID Data Base



 $\gamma p \rightarrow \pi^0 p$: Bump structure near W=1.65-1.7 GeV

: Almost no data

There is a need to study $\gamma n \rightarrow \pi^0 n$

 π^0 Data analysis

- Detection of two photons in the BGO ball and the recoil proton/neutron in the forward detectors;

- Cut on the neutron/proton missing mass;
- Cut on the 2γ invariant mass;
- Cut on the square of the pion missing mass;
- Complanarity;



Reaction yields



N

 $\gamma p \rightarrow \pi^0 p$: Bump structure at W~1.7 GeV

γn→π^on: No bump structure

See also Talk of Bernd Krusche by Monday.

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Rejection of $\pi 0$ background



Symmetric decays are rejected by the analysis of cluster shapes in the BGO Ball. Efficiency of this rejection is ~99%. If the pion is emitted at backward angles, its energy is low. Such events are suppressed.



Asymmetric decays: If the first photon is emitted at **the backward angles**, the low-energy second photon can be detected in the Russian Wall or the BGO Ball.

 \rightarrow Discrimination of Compton scattering from π^{0} events is possible at backward angles.

Data analysis

- Detection of only one photon in the BGO Ball and the recoil proton/neutron in the Russian Wall;

- Complanarity;
- Cut in the neutron/proton TOF and angle;
- Cut on the proton/neutron missing mass;
- Cut on the squared photon missing mass;
- Cluster-shape analysis.

After that Only Compton and π^0 events (asymmetric decays) are selected.



$\gamma p \rightarrow \gamma p$ on the free proton at 150< Θ cm<165 deg

Data

Simulations





Cross Check



Different Binning





Summary

First results on Compton scattering on the neutron reveal the peak at $W \approx 1.685$ GeV which is not seen in Compton scattering on the proton. Being considered in conjunction with the results on eta photoproduction, this observation supports the existence of N*(1685).

Thanks for your attention!

 π^0 Data analysis

- Detection of two photons in the BGO ball and the recoil proton/neutron in the forward detectors;

- Cut on the neutron/proton missing mass;
- Cut on the 2y invariant mass;
- Cut on the square of the pion missing mass;
- Complanarity;



Reaction yields



 $\gamma n \rightarrow \pi^0 n$: No bump structure

 $\gamma p \rightarrow \pi^0 p$: Bump structure at W~1.7 GeV

See also Talk of Bernd Krusche by Monday.

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